



Crystalline Coatings

a new paradigm in optical coating technology

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Faculty of Physics, University of Vienna, Austria

Crystalline Mirror Solutions GmbH & LLC



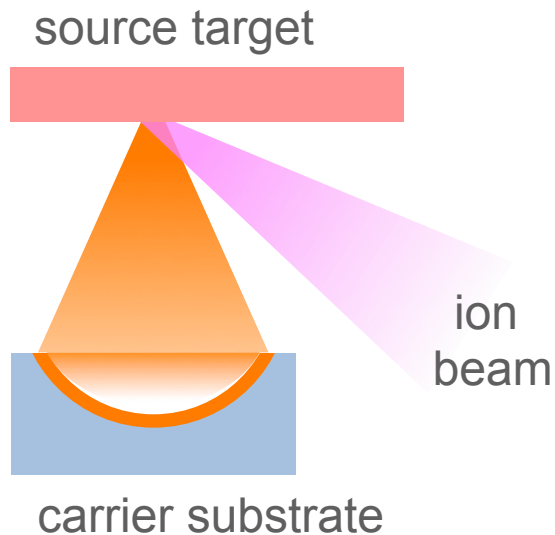
Redefining precision laser optics



universität
wien

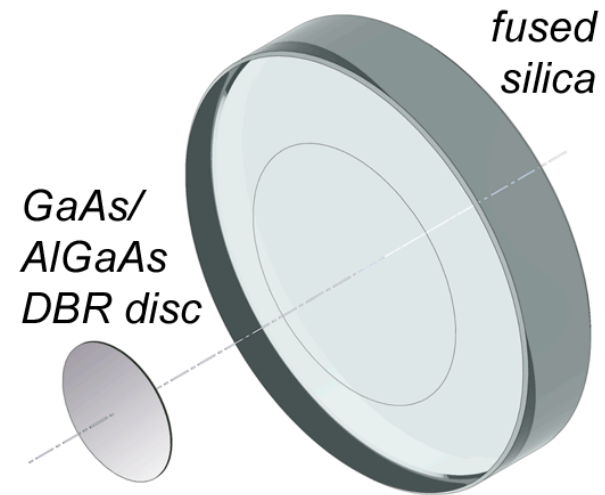
- Introduction to CMS
- NIR noise- and optical performance
 - demonstration of sub-ppm absorption and <5 ppm scatter
 - demonstration of ultra-low coating thermal noise
- MIR optical performance ($> 2\mu\text{m}$)
 - demonstration of sub-100ppm absorption losses @ 4mm
 - potential for significantly lower optical loss-levels
- Application Examples

State-of-the-art technology – sputtered dielectric mirrors



-30 years of in-market experience

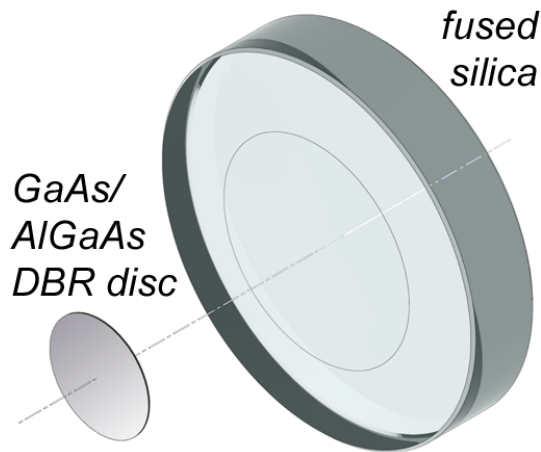
New technology – **substrate transferred crystalline coatings**



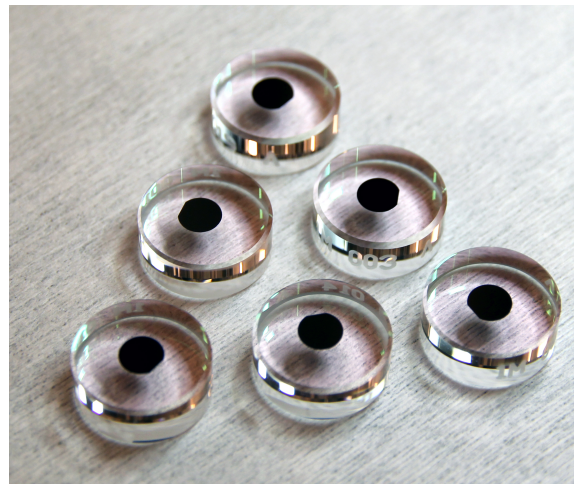
<3 years of in-market experience

... a spin-off of fundamental research from the University of Vienna

- ❖ We offer an entirely unique technology for precision laser optics
 - Applications span fundamental R&D + emerging industrial uses
 - cutting-edge research efforts in spectroscopy and metrology
 - advanced industrial applications in manufacturing and sensing



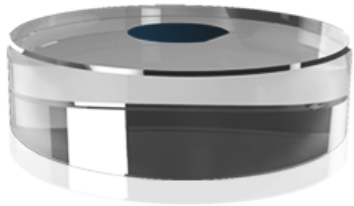
**substrate-transferred
single-crystal coatings**



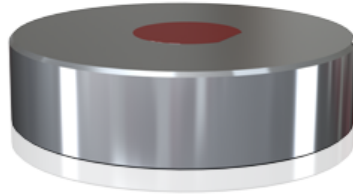
**end mirrors for
ultrastable cavities**



**cavity employed for
laser stabilization**



◆ xtal stable




◆ xtal mir



◆ xtal therm

- We are offering **three distinct product lines** based on the unique performance advantages of semiconductor crystalline coatings
 - **xtal stable:** Ultralow Brownian noise for cavity-stabilized lasers used in optical metrology, laser ranging, and inertial navigation systems
 - **xtal mir:** Low optical losses in the mid-IR (projected loss <100 ppm to $\sim 5 \mu\text{m}$ and beyond) for spectroscopy / trace gas analysis
 - **xtal therm:** High thermal conductivity (50-100 \times greater than amorphous films) reflectors for high power (i.e. thin disk) lasers



BREAKING NEWS – Jan 11 2016, 5pm CET

„We have detected a gravitational wave. It is from two black holes with masses in the 30 solar mass range and is a picture perfect fit to what Einstein predicts. „

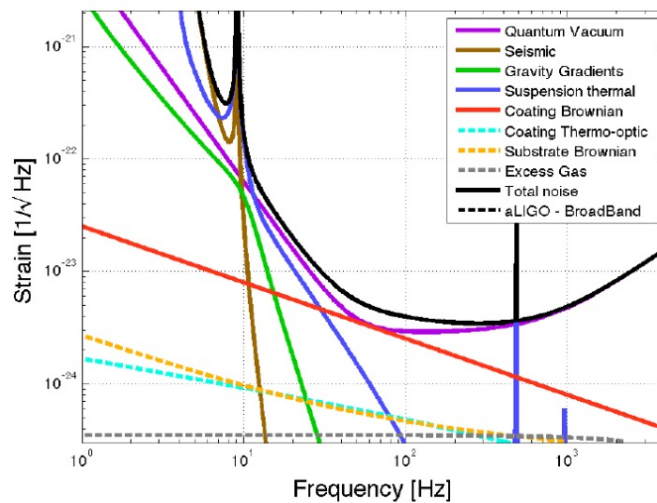
CONGRATULATIONS LIGO!!

LIGO – Laser Interferometer Gravitational Wave Observatory

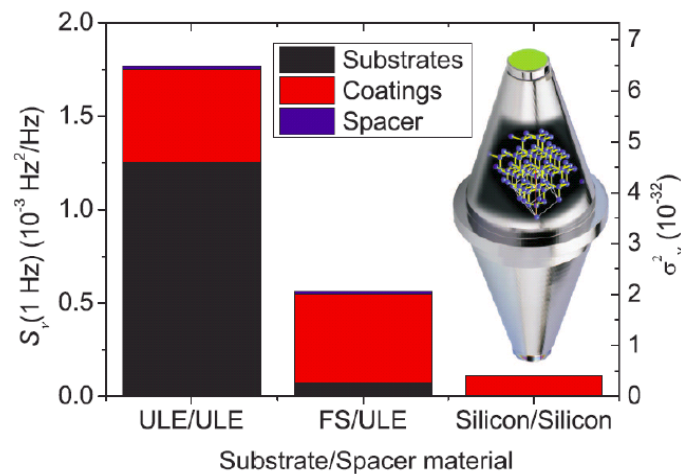
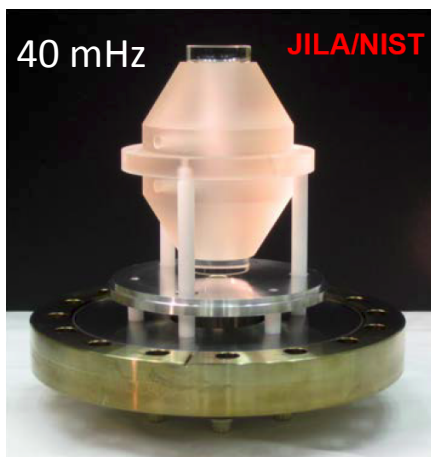
Gravitational wave detectors



Advanced LIGO



Ultrastable lasers



Coating thermal noise

Thermal noise of optical coatings: fundamental limit for laser frequency stabilization with cavities (atomic clocks, precision spectroscopy, gravitational wave detectors, ...)

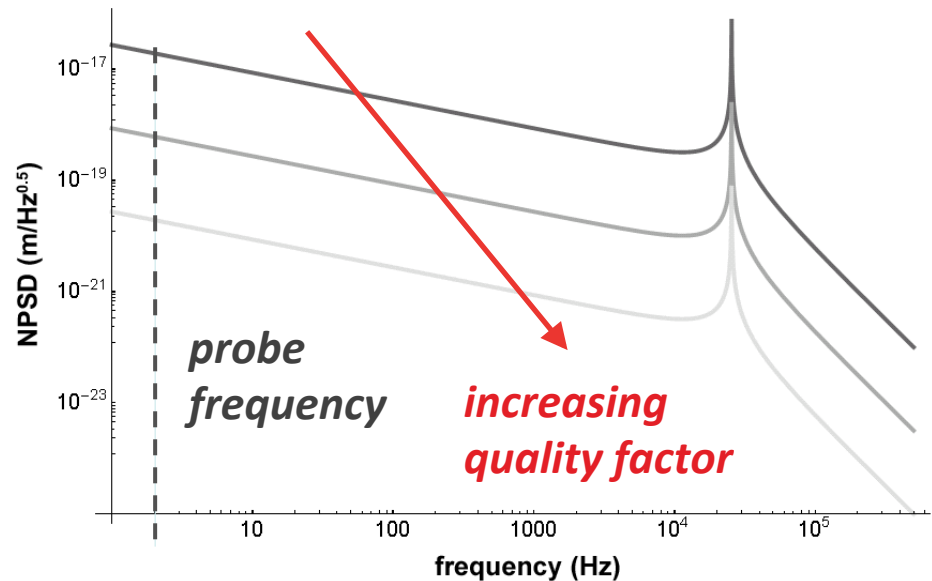
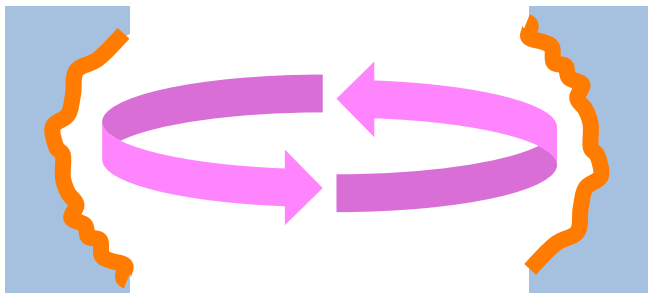
Penn et al., Class. Quant. Grav. 20, 2917 (2003)
Numata et al., PRL 93, 250602 (2004)
Notcutt et al., PRA 73, 31804 (2006)
Kessler et al., JOSA B 29, 117 (2012)
 ...

Fluctuation–Dissipation Theorem:

$$\left[\delta x(\omega) = \left(\frac{2k_B T}{\pi m} \gamma \right)^{1/2} |X_{\text{mech}}| \right] \propto \sqrt{\gamma} = \sqrt{\frac{1}{Q}} = \sqrt{\Phi}$$

δx : Thermally driven displacement fluctuations \rightarrow cavity phase noise

Φ : Loss angle



Thermal noise of optical coatings: fundamental limit for laser frequency stabilization with cavities (atomic clocks, precision spectroscopy, gravitational wave detectors, ...)

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Monocrystalline optomechanics:
 high optical AND mechanical quality
 \rightarrow low coating thermal noise

AlGaAs crystalline coatings

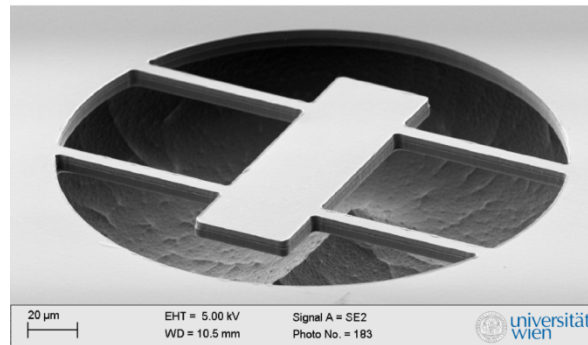
$\Phi_{\text{coat}} \sim 2 \times 10^{-5}$ @ 300K

$\Phi_{\text{coat}} \sim 5 \times 10^{-6} - 1 \times 10^{-5}$ @ 4K

vs.

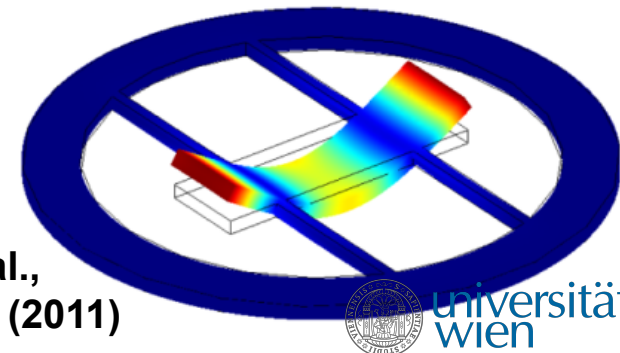
SiO₂/Ta₂O₅ coatings

$\Phi_{\text{coat}} > 4 \times 10^{-4}$ @ 300K



Cole et al., APL 92, 261108 (2008)

Cole, Wilson-Rae et al., Nature Comm. 2, 231 (2011)



Thermal noise of optical coatings: fundamental limit for laser frequency stabilization with cavities (atomic clocks, precision spectroscopy, gravitational wave detectors, ...)

Penn et al., *Class. Quant. Grav.* 20, 2917 (2003)
 Numata et al., *PRL* 93, 250602 (2004)
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AlGaAs crystalline coatings

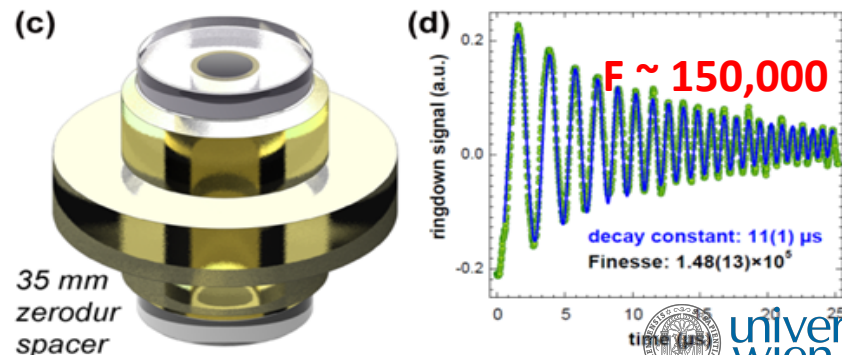
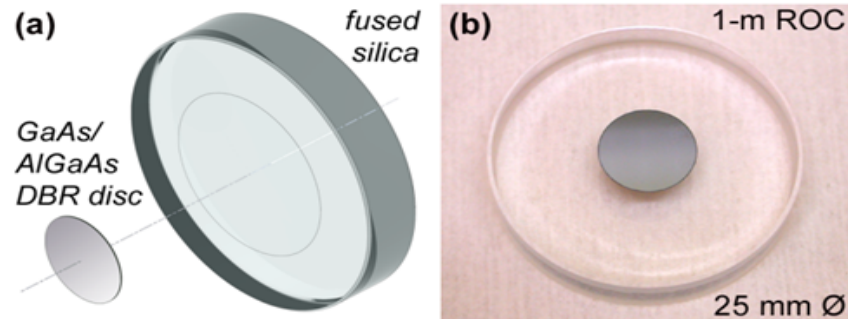
$$\Phi_{\text{coat}} \sim 2 \times 10^{-5} \text{ @ } 300\text{K}$$

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vs.

SiO₂/Ta₂O₅ coatings

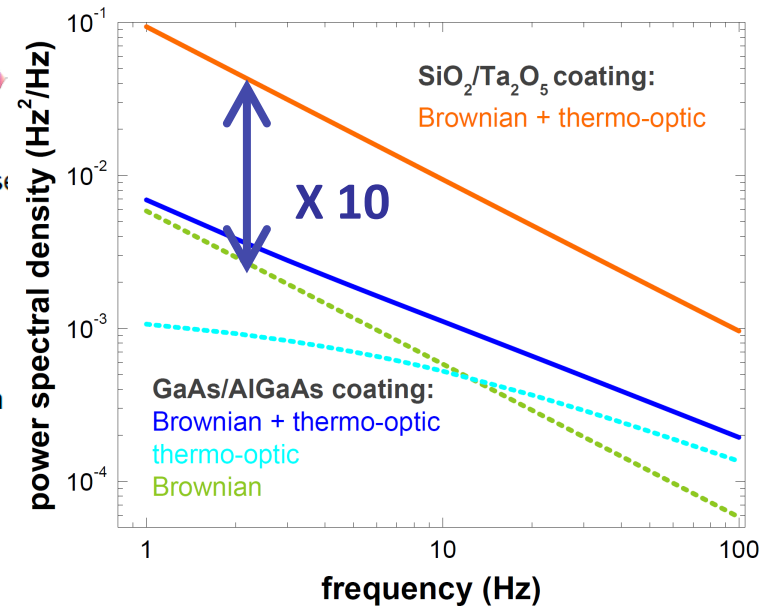
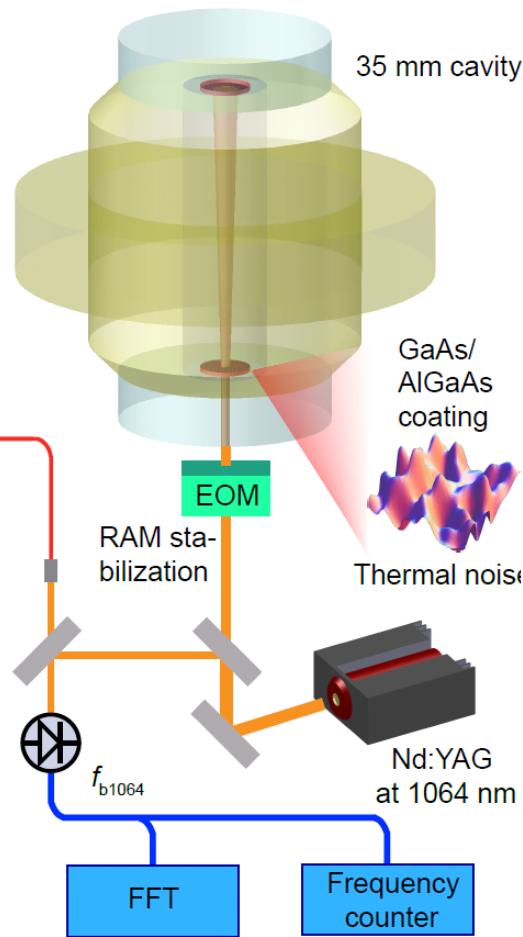
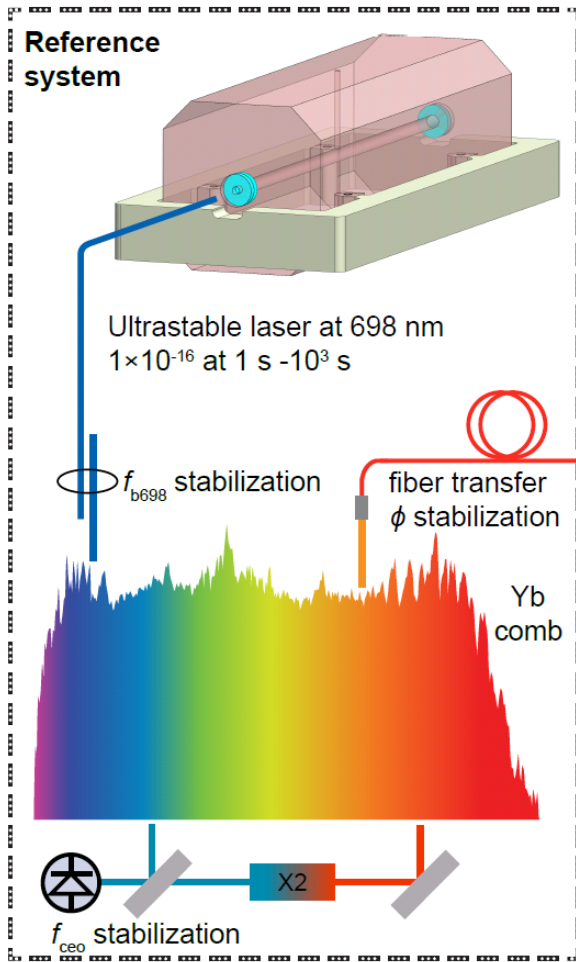
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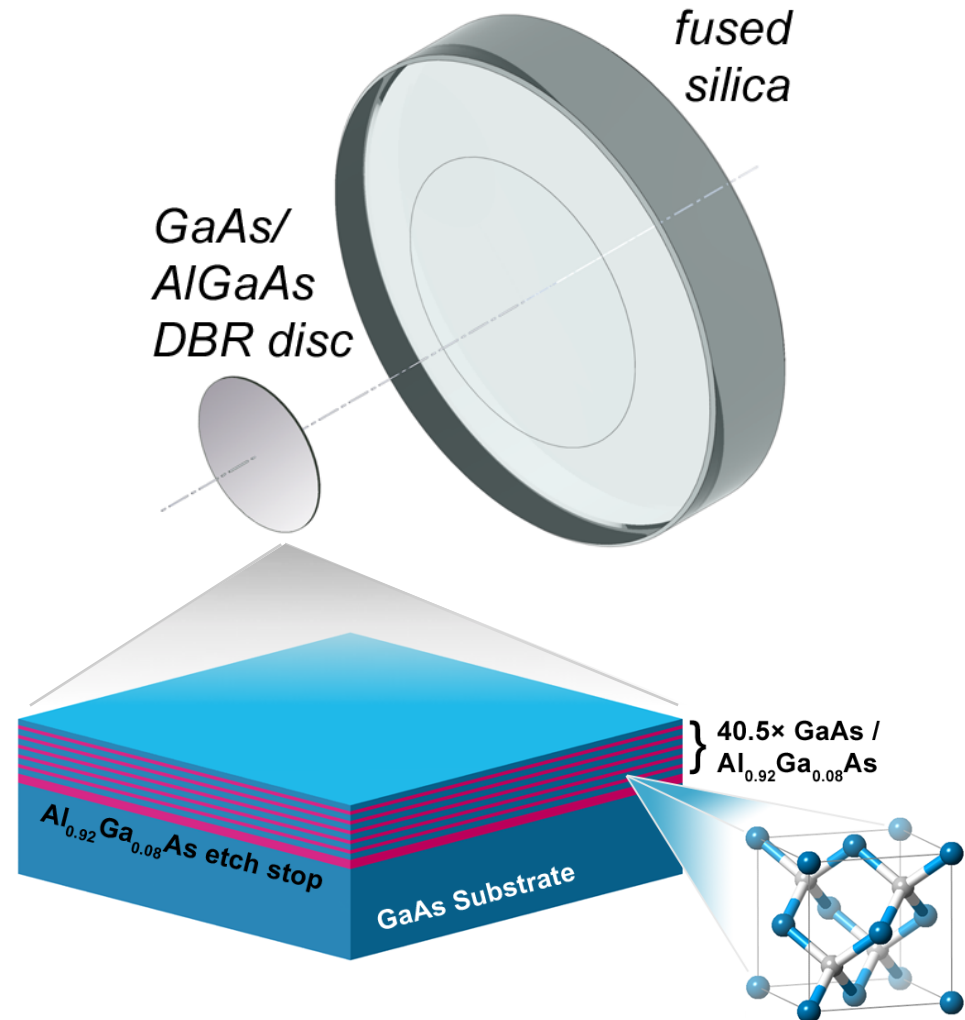
Optomechanics of optical coatings

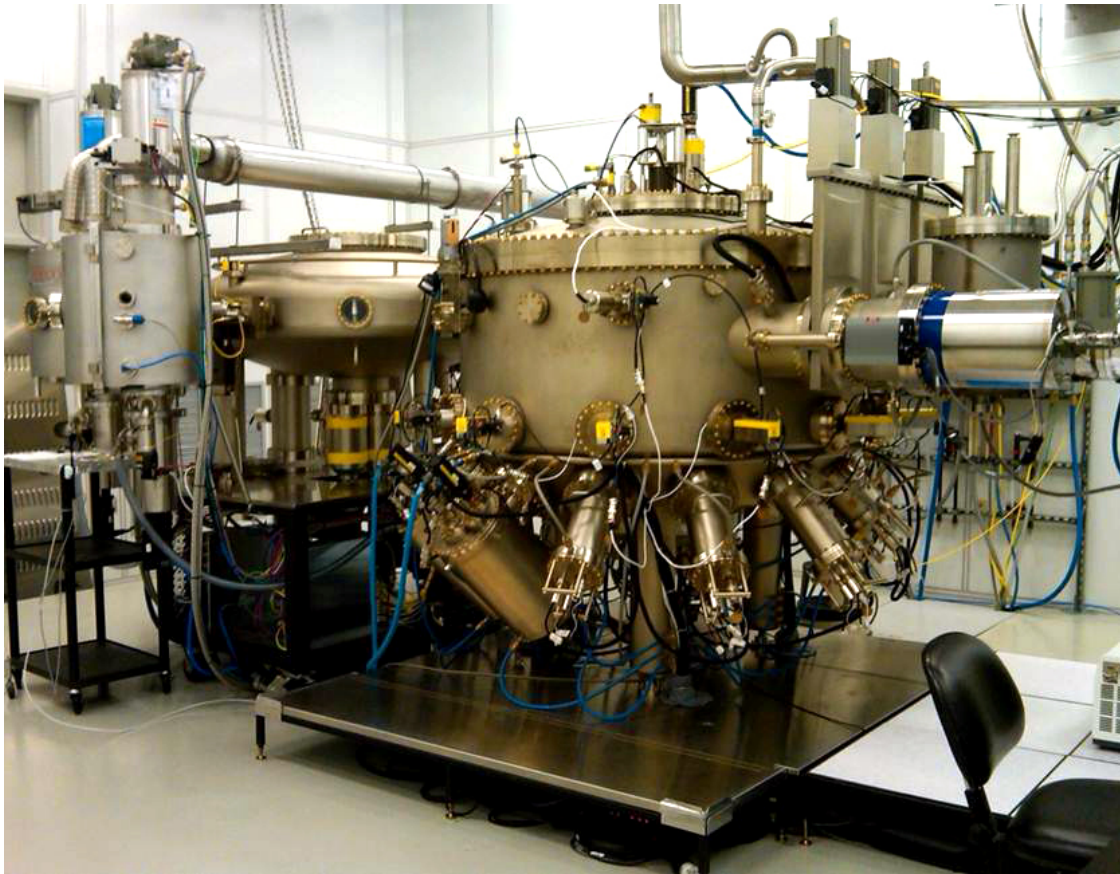
aLIGO

CMS | **CRYSTALLINE MIRROR SOLUTIONS**



- AlGaAs multilayer with varying Al content for index contrast
 - high index layers consist of binary GaAs thin films
 - 8% Ga incorporated in low index AlGaAs layers to slow oxidation in ambient
- Potential for high reflectivity from ~ 900 nm \rightarrow $5+$ μ m
 - peak performance in NIR
- High quality epitaxy requires a lattice matched substrate
 - same crystalline symmetry
 - minimal deviation of lattice parameter (atomic spacing)

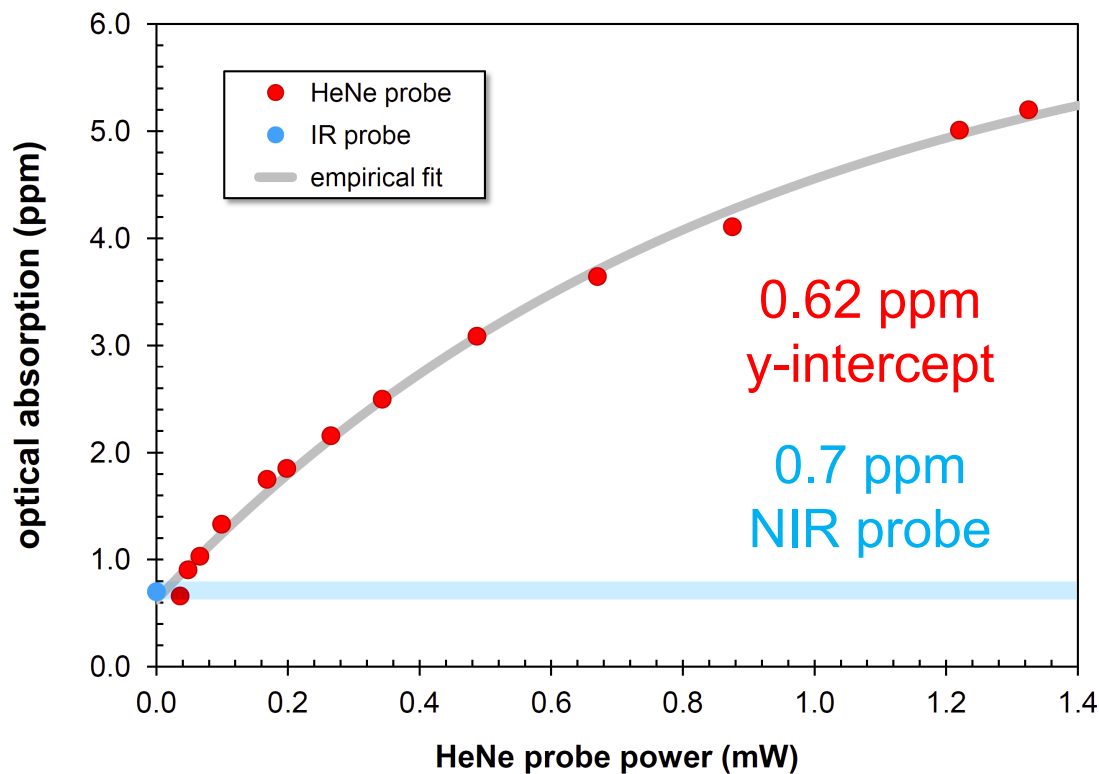




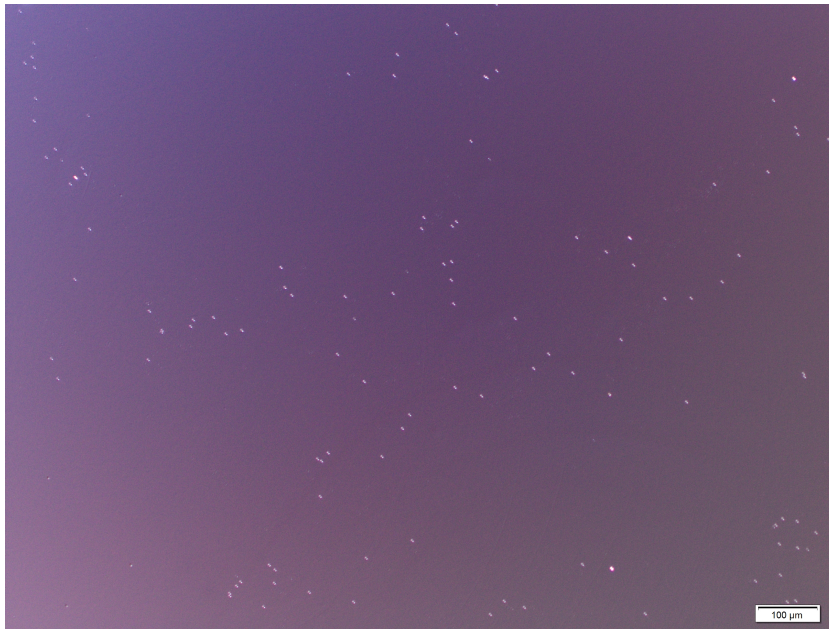
Since June 2016:
in-house MBE growth (tool operated on IBM Research campus in Zurich)



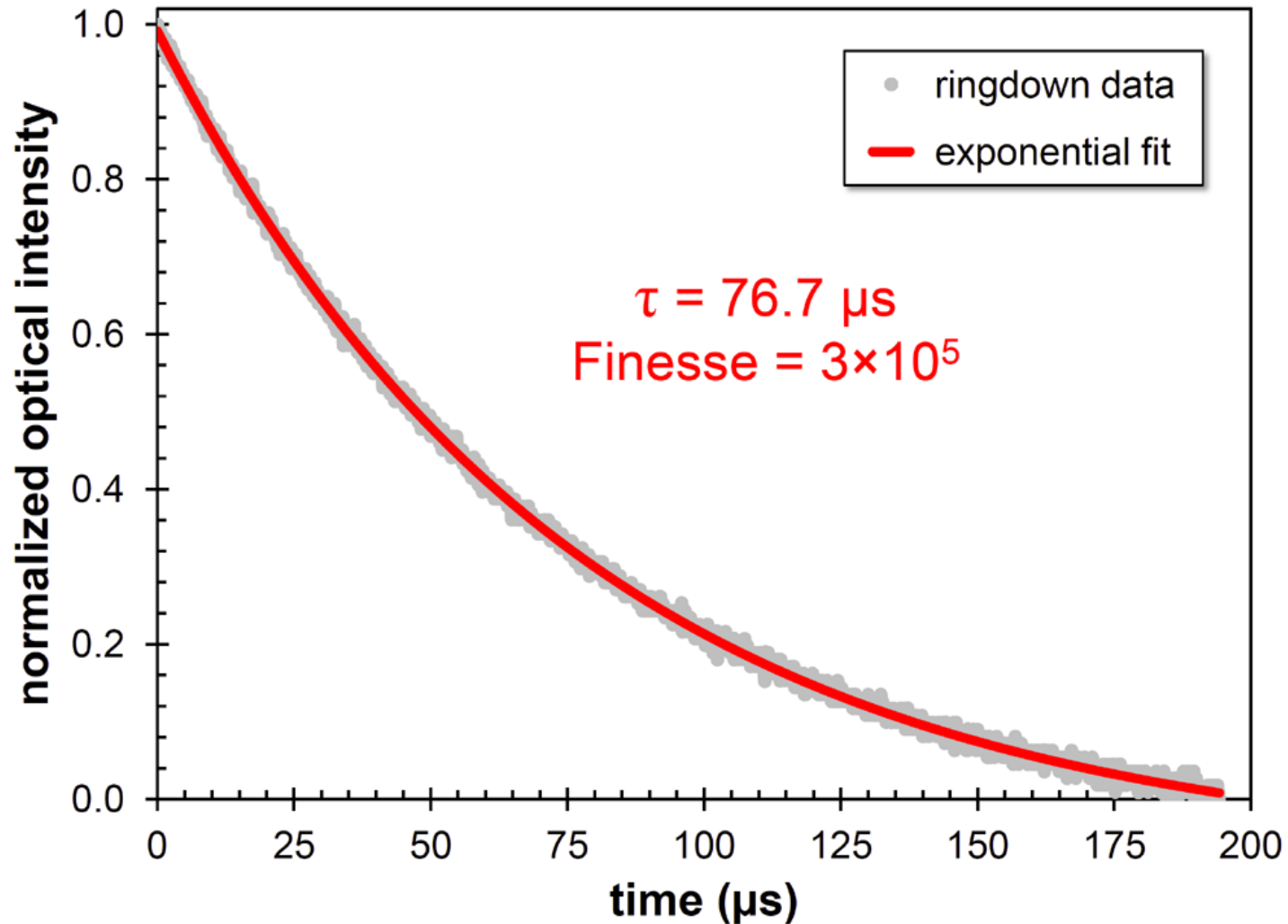
- Absorption study run with our epitaxial growth foundry
- Samples sent to SPTS, Pahoia, HI for absorption measurements



- Red HeNe probe in PCI system limits minimum absorption value
 - the use of a transparent near-IR probe is now being investigated
- Probe-power dependent measurements yield sub-ppm absorption



- Embedded “oval defects” incorporated during MBE growth process
 - leads to strong position dependence in optical loss of coatings
- Defect free regions have shown micro-roughness-limited losses
 - excess losses (scatter + absorption) verified at the <5 ppm level



Substrate-transferred crystalline coatings simultaneously exhibit **excellent optical and mechanical quality**

- ◆ Damping reduction of **10-100×** compared with IBS films
 - IBS-deposited $\text{Ta}_2\text{O}_5/\text{SiO}_2$: typical $Q \sim 3000$ ($\phi_{\text{IBS}} \approx 2-4 \times 10^{-4}$)
 - AlGaAs room temperature Q-value of **4×10^4** ($\phi_{\text{RT}} \approx 2 \times 10^{-5}$)
 - AlGaAs cryogenic performance: $Q > 1 \times 10^5$ ($\phi_{\text{min}} \approx 4.5 \times 10^{-6}$)
- ◆ Minimal scattering loss and optical absorption
 - absorption verified at **< 1 ppm** (**$< 0.01 \text{ cm}^{-1}$** at 1064 nm)
 - RMS micro-roughness of **1.3 \AA RMS** (**< 3 ppm** at 1064 nm)
- ◆ Excess losses (S+A) down to levels **< 5 ppm**
 - measured finesse of **$> 3 \times 10^5$** at 1064, 1397 and 1550 nm

Multiple order-of-magnitude improvements

❖ xtal stable

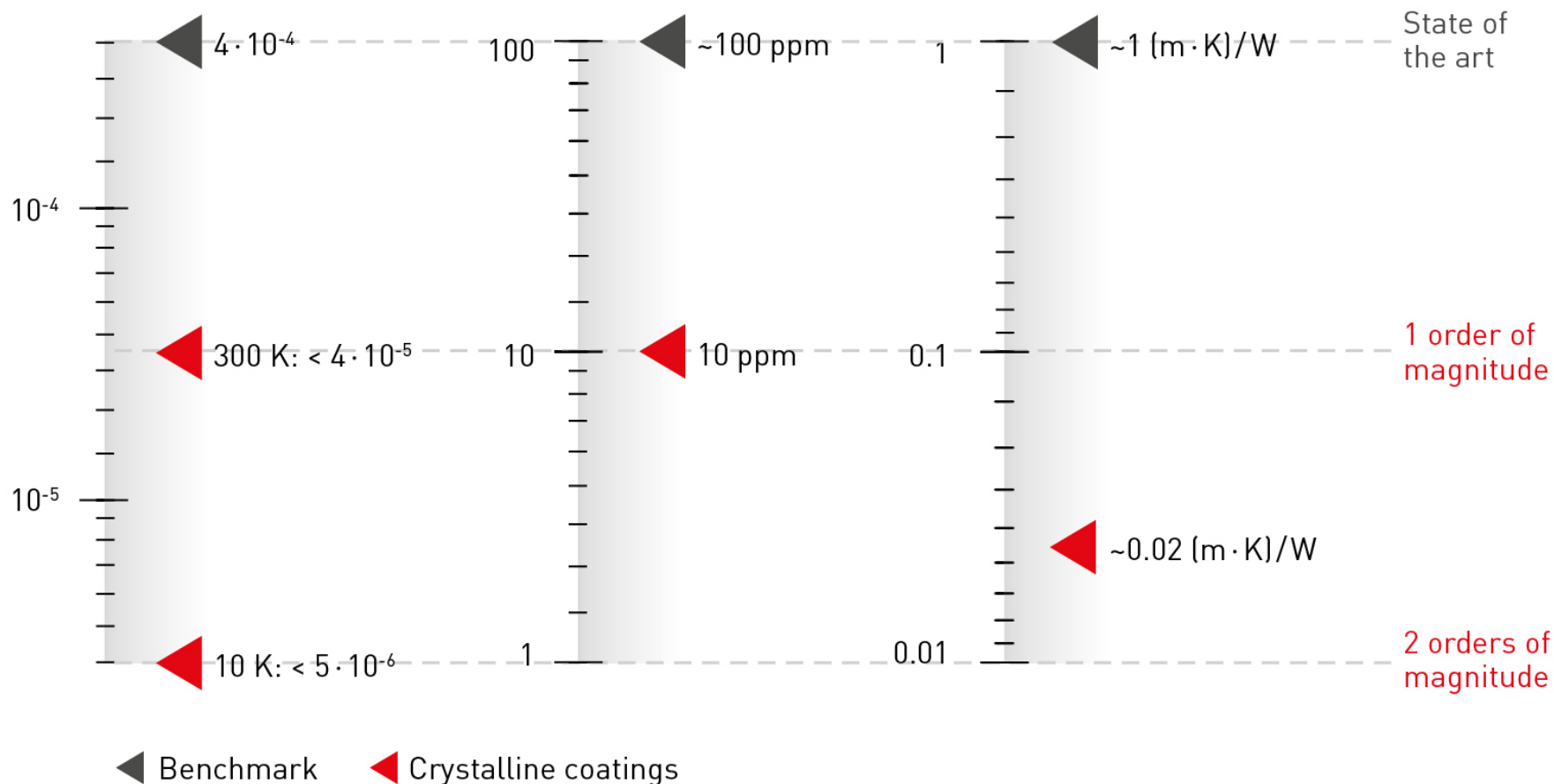
Reduced Brownian noise
(mechanical loss angle)

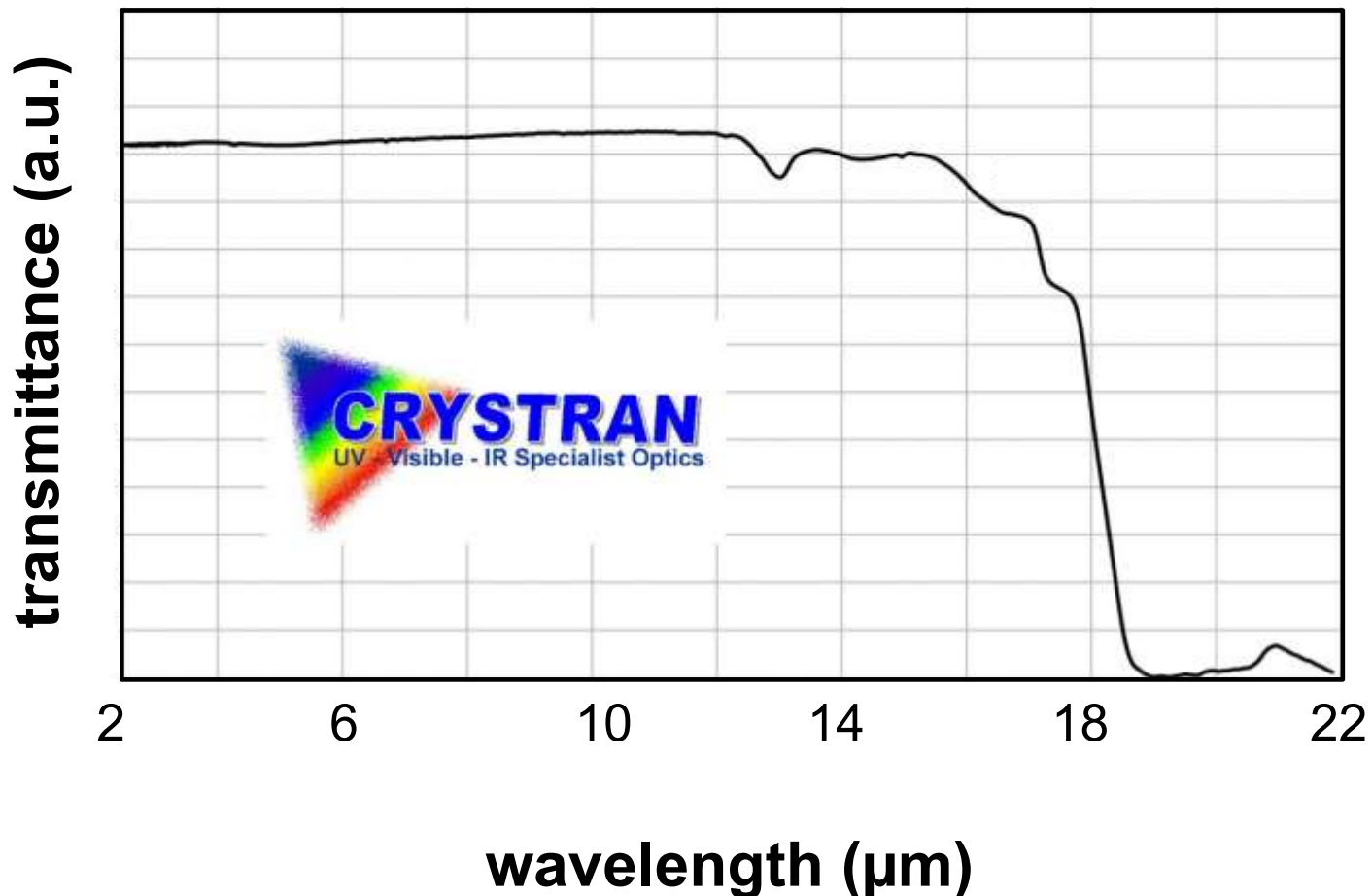
❖ xtal mir

Reduced mid-IR
optical absorption

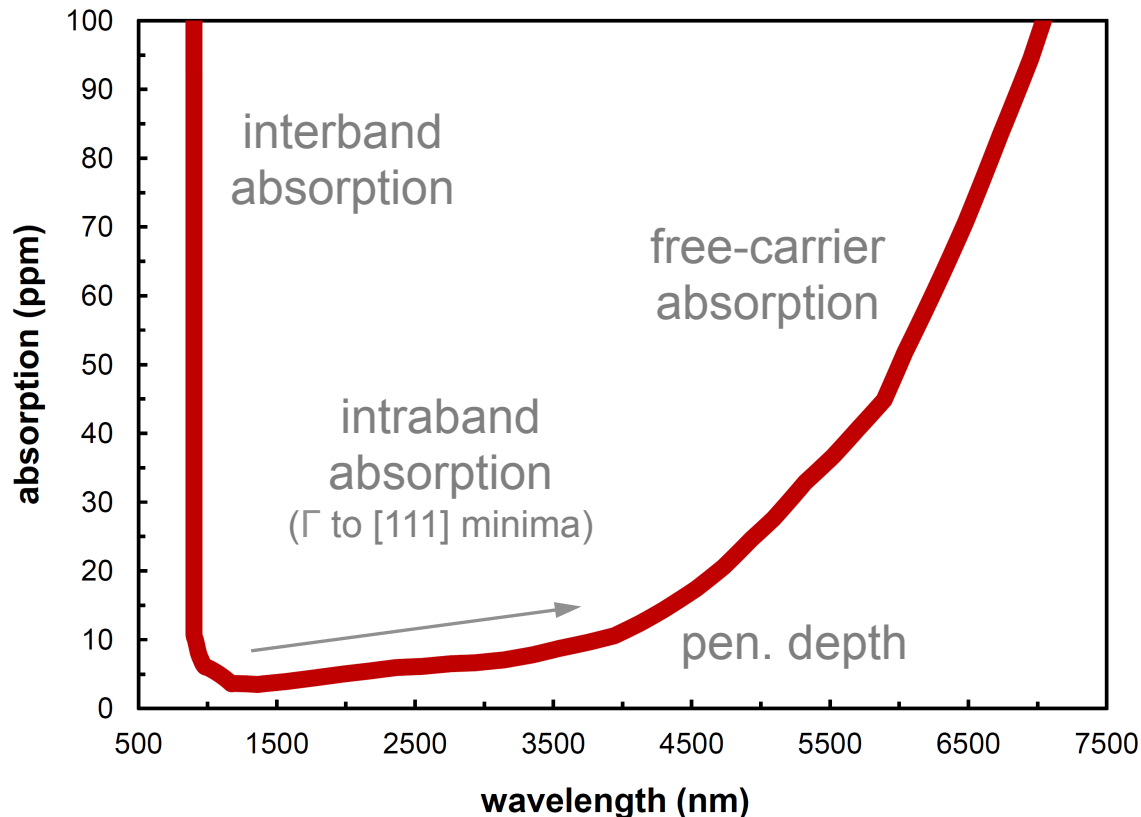
❖ xtal therm

Reduced thermal
resistivity





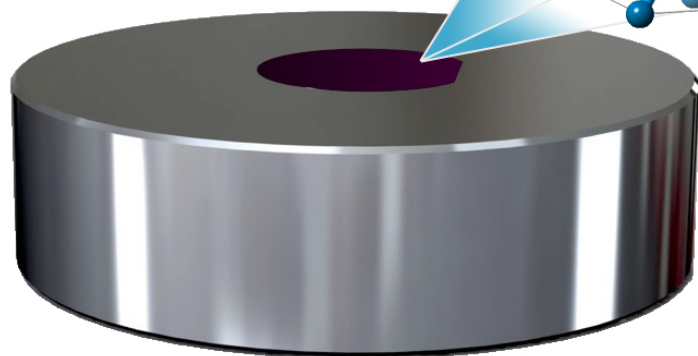
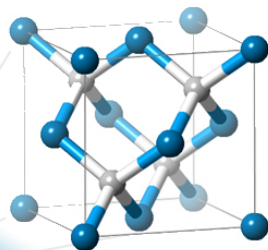
GaAs is a commonly used material for mid and long-wave optics
What losses can be expected from GaAs/AlGaAs multilayers?



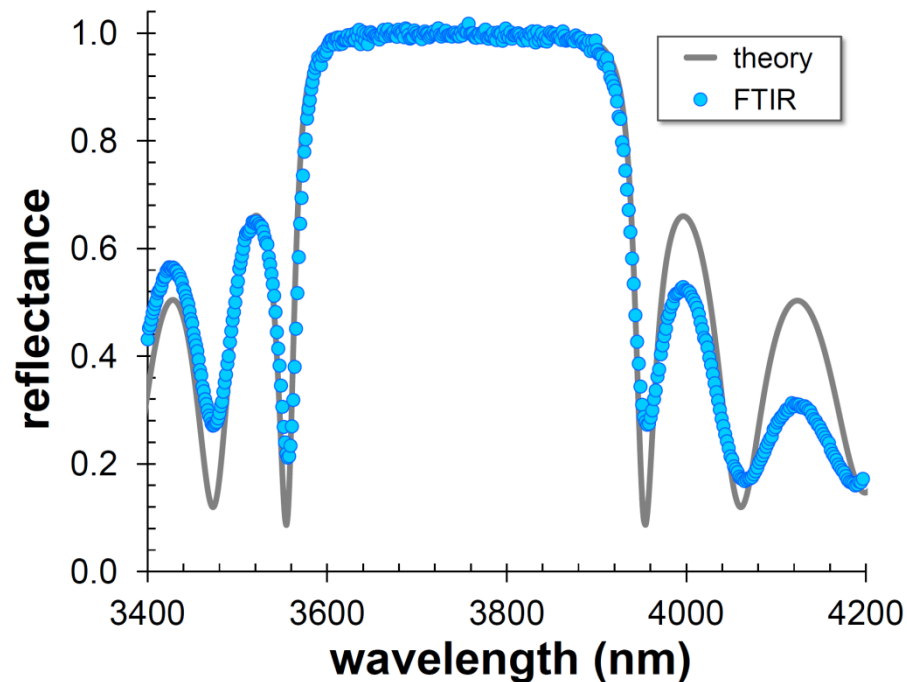
- Plot incorporates
 - material dispersion
 - absorption coefficient
 - penetration depth
- Assumes n-type background doping
- GaAs band-edge limits operation to >900 nm
 - short-wave cutoff
- Long-wavelength cutoff beyond $5 \mu\text{m}$
 - free-carrier absorption
 - proportional to λ^3

For abs. coefficient see: W. G. Spitzer and J. M. Whelan, Physical Review 114, April 1959

28.5 period
GaAs/AlGaAs
crystalline coating

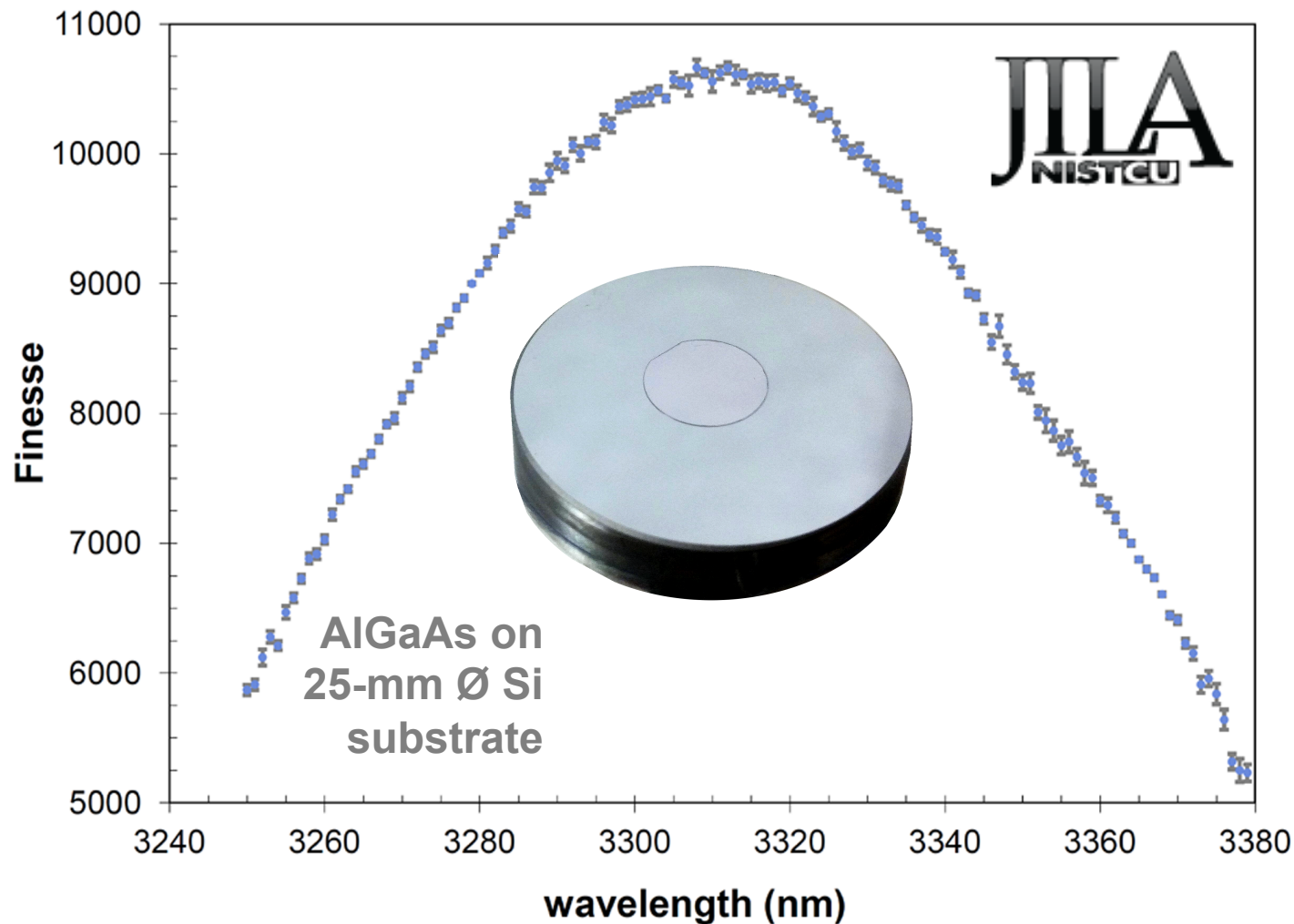


single-crystal Si (SCS) substrate



- Monocrystalline mirror discs transferred to curved SCS substrates
 - 28.5 period GaAs/Al_{0.92}Ga_{0.08}As DBR, 600 ppm trans. @ 3725 nm
 - potential for optical losses below 100 ppm up to ~5 μ m

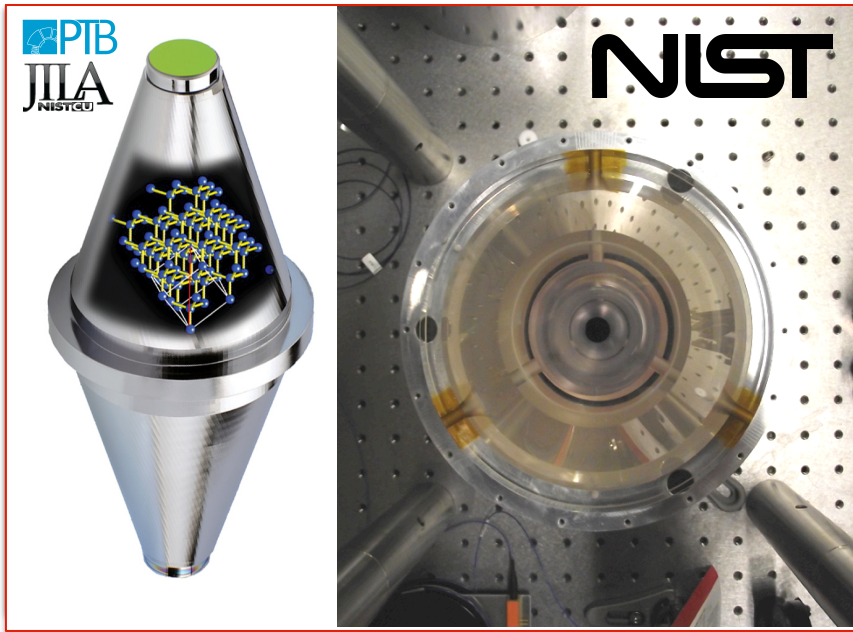
Preliminary MIR Mirror Results



Collaboration with B. Bjork, O. Heckl, B. Spaun, B. Changala, J. Ye, JILA, Boulder, CO

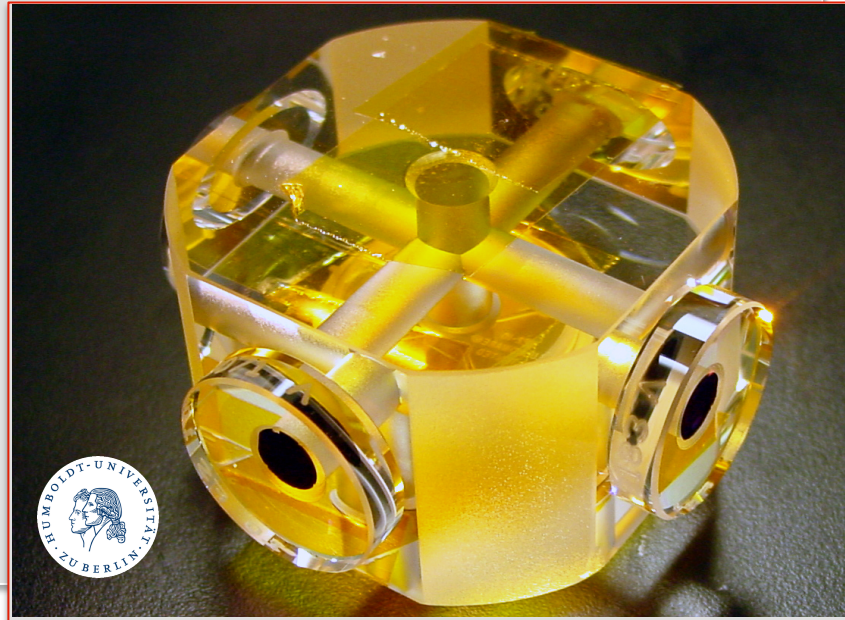
towards frequency
stability $\Delta f/f$ of 1×10^{-17}
and beyond

optical atomic clocks



Zhang et al., Opt. Lett. 39, 1980 (2014)

Cole et al., Optica 3, 647 (2016)



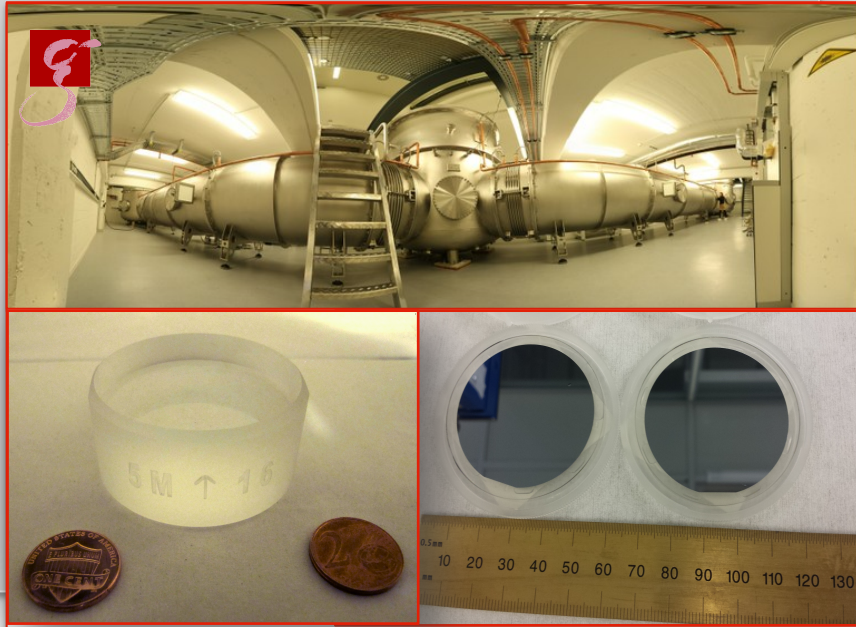
optical atomic clocks

test of fundamental
symmetries

• optical atomic clocks

• test of fundamental symmetries

• gravitational wave detector optics



Chalermongsak et al., Metrologia 53, 860 (2016)

J. Steinlechner et al., CQG 32, 105008 (2015)

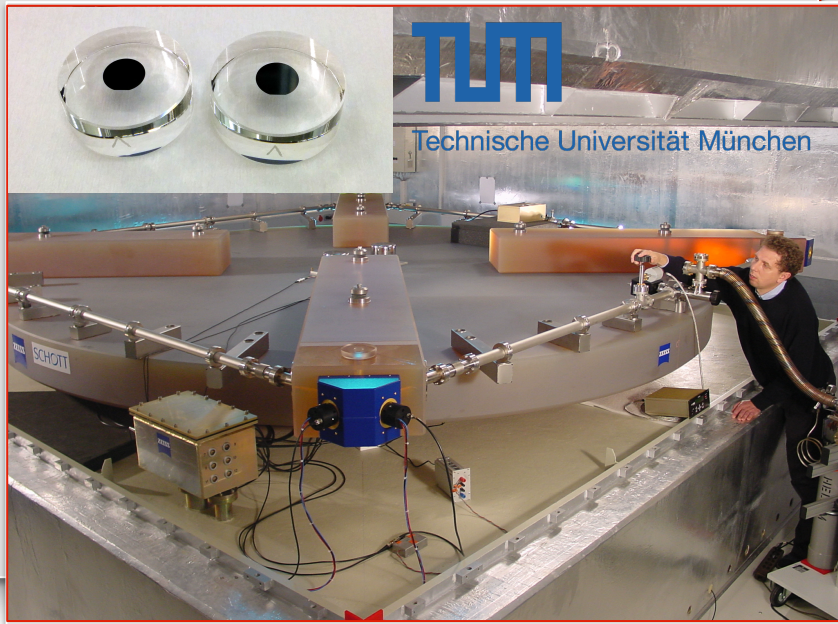
2.56 m² helium–neon based ring laser
interferometer at a wavelength of 1.152 μm

optical atomic clocks

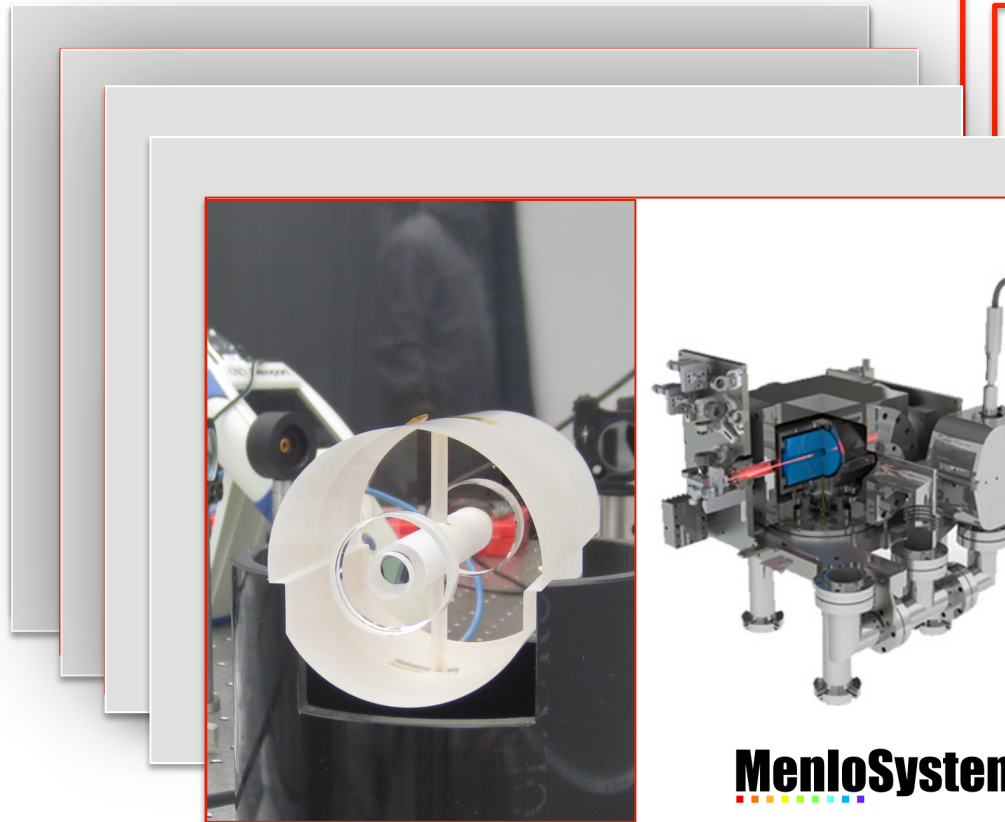
test of fundamental
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gravitational wave
detector optics

ring-laser gyroscopes



Schreiber et al., Opt. Lett. 40, 1705 (2015)



optical atomic clocks

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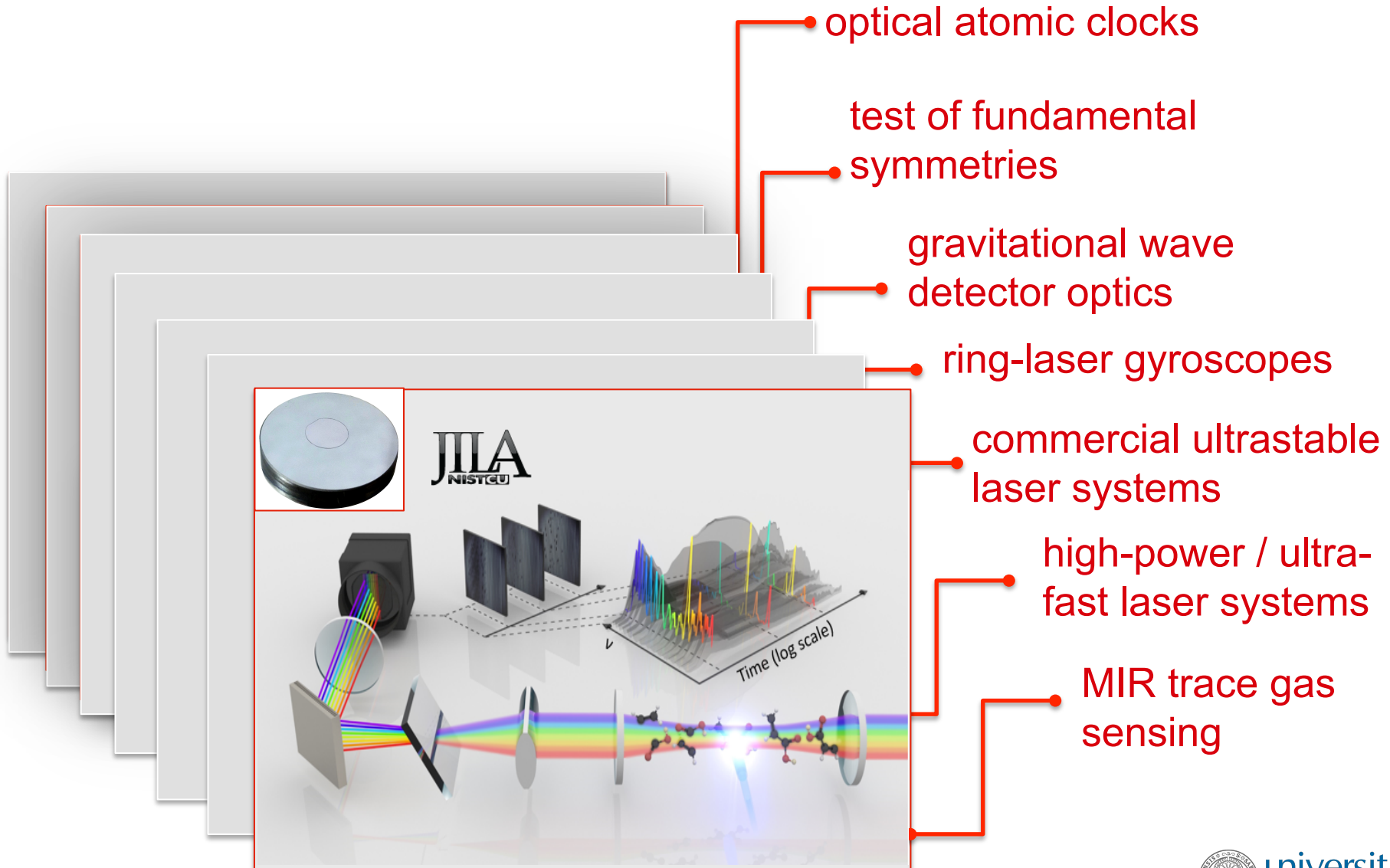
ring-laser gyroscopes

commercial ultrastable laser systems

- GaAs-based SESAMs transferred to SiC
- superior flatness to competing technologies

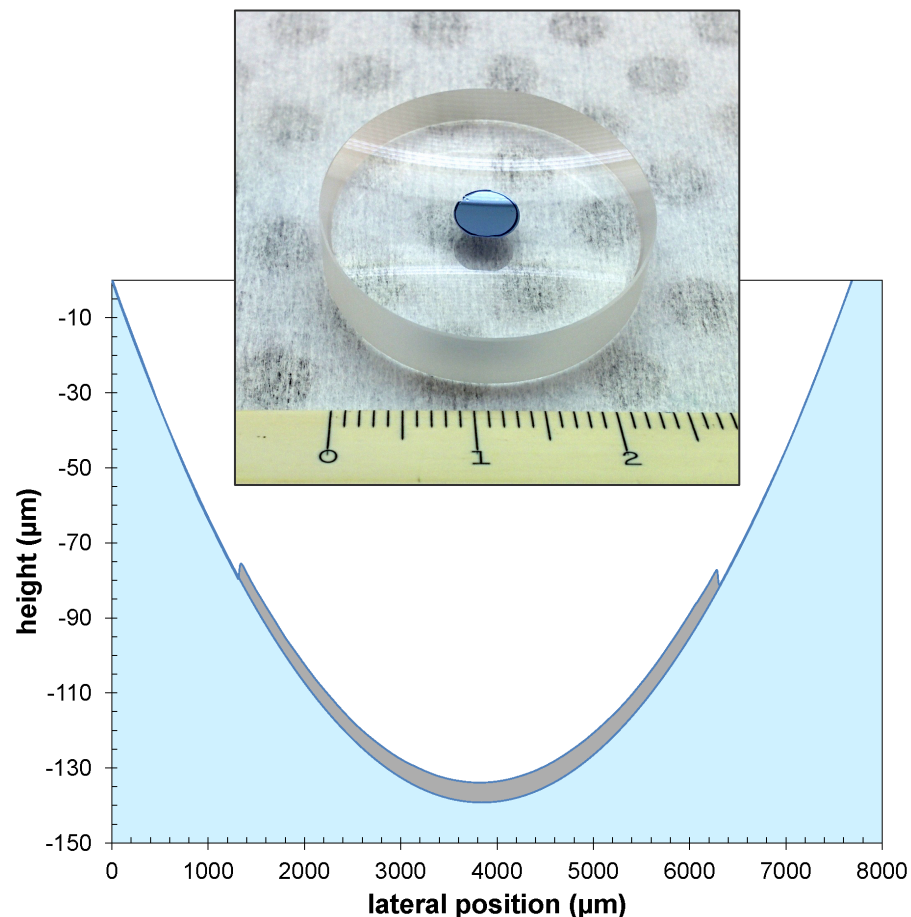


- optical atomic clocks
- test of fundamental symmetries
- gravitational wave detector optics
- ring-laser gyroscopes
- commercial ultrastable laser systems
- high-power / ultra-fast laser systems



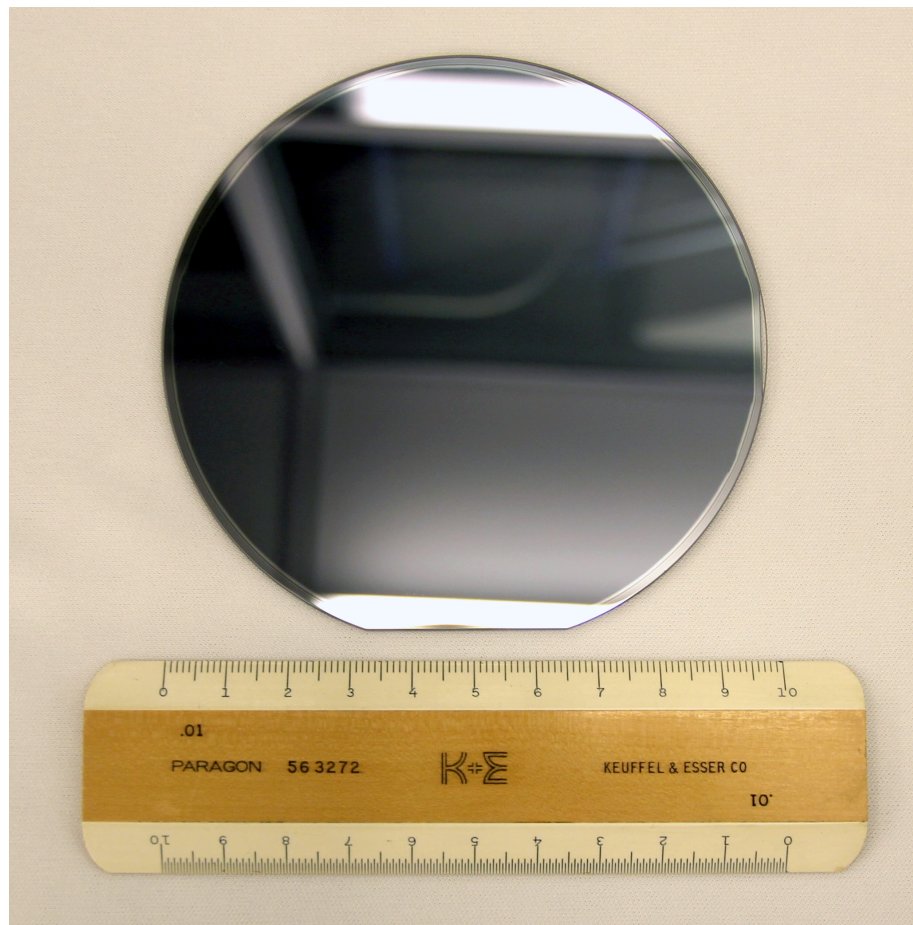
B. Bjork, J. Ye et al., submitted (2016)

- High optical quality in the mid-infrared (2-5 μm)
 - initial tests with Jun Ye's group reveal scatter + absorption losses at the ~ 100 ppm level (3725 nm)
- High thermal conductivity
 - ~ 70 W/m \cdot K (IBS: ~ 1)
- Variety of substrate options
 - Si, SiC, diamond, sapphire
- Transfer to curved surfaces
 - minimum ROC: 100 mm
 - larger ROC is easier
- 100+ mm \varnothing direct bonding
 - void free over entire area
 - bond strengths of ~ 1 J/m 2



Transfer to curved surfaces:
minimum achievable ROC of 10 cm

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Large area direct bonding:
100-mm diameter GaAs on silica

Semiconductor-based substrate-transferred crystalline coatings - pushing the limits of ultimate laser performance

> 10-fold improvement in...

...Brownian noise (for optical clocks, gravitational wave detectors, ...)

Cole et al., Nature Phot. 7, 664 (2013)

Zhang et al., Opt. Lett. 39, 1980 (2014)

Schreiber et al., Opt. Lett. 40, 1705 (2015)

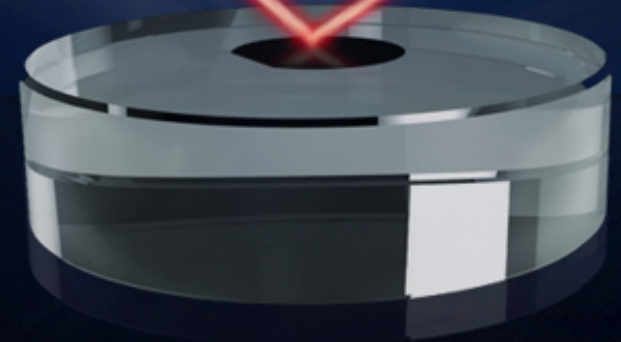
Chalermsongsak et al., Metrologia 53, 860 (2016)

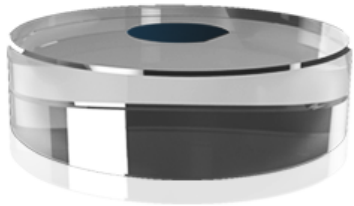
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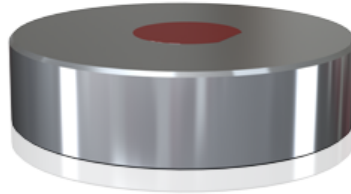
...thermal conductivity (for high-power lasers, SESAMs, etc.)

Diebold et al., Opt. Expr. 24, 10512 (2016)





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Thanks to the CMS team



Thank You For Your Attention!

